

**Ускорительные  
нейтринные  
эксперименты:  
последние  
новости**

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# Mixing matrix

Pontecorvo-Maki-Nakagawa-Sakata matrix

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = U \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

atmospheric

solar

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix} \begin{pmatrix} \cos\theta_{13} & 0 & \sin\theta_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin\theta_{13}e^{-i\delta} & 0 & \cos\theta_{13} \end{pmatrix} \begin{pmatrix} \cos\theta_{12} & \sin\theta_{12} & 0 \\ -\sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

link between atmospheric and solar

Oscillation experiments



$\Delta m_{12}^2$   $\Delta m_{13}^2$   
 $\sin\theta_{12}$   $\sin\theta_{23}$   $\sin\theta_{13}$

$\delta$  - CP violating phase

# Neutrino mass and mixings

- 3 mixing angles ( $\theta_{12}, \theta_{23}, \theta_{13}$ )
- 1 CPV phase ( $\delta$ )
- 2 (independent) mass differences ( $\Delta m_{ij}^2 = m_i^2 - m_j^2$ )

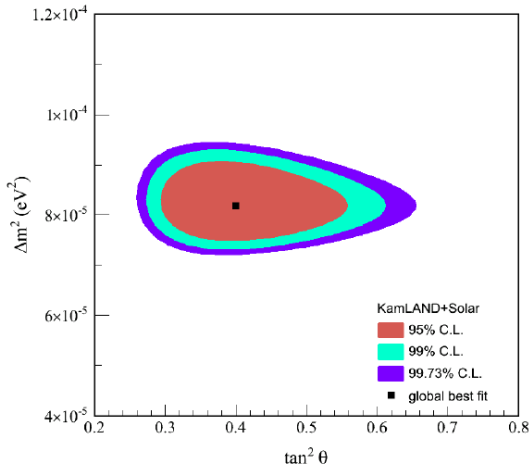
$$\theta_{12}, \Delta m_{12}^2$$

$$\theta_{23}, \Delta m_{32}^2$$

$$\theta_{13}, \Delta m_{31}^2$$

$$\Delta m_{\text{solar}}^2 = 8 \times 10^{-5} \text{ eV}^2$$

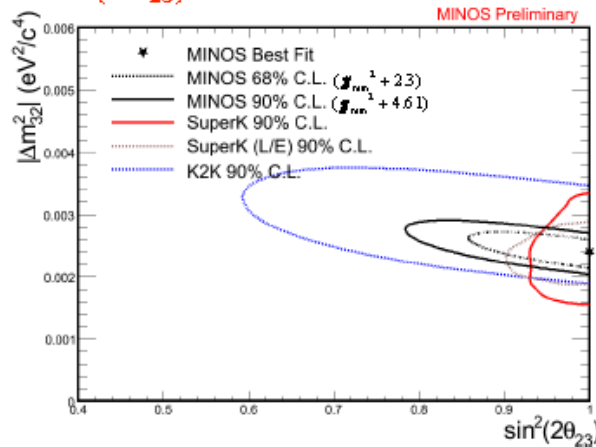
$$\sin^2(2\theta_{12}) = 0.86$$



Solar + KamLAND

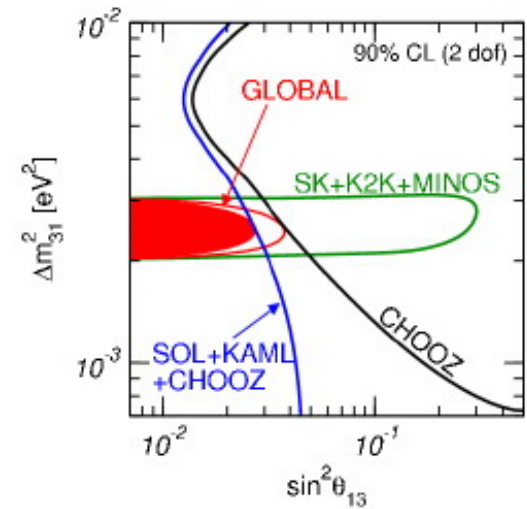
$$\Delta m_{\text{atm}}^2 = (2.2 \sim 2.6) \times 10^{-3} \text{ eV}^2$$

$$\sin^2(2\theta_{23}) > 0.92$$



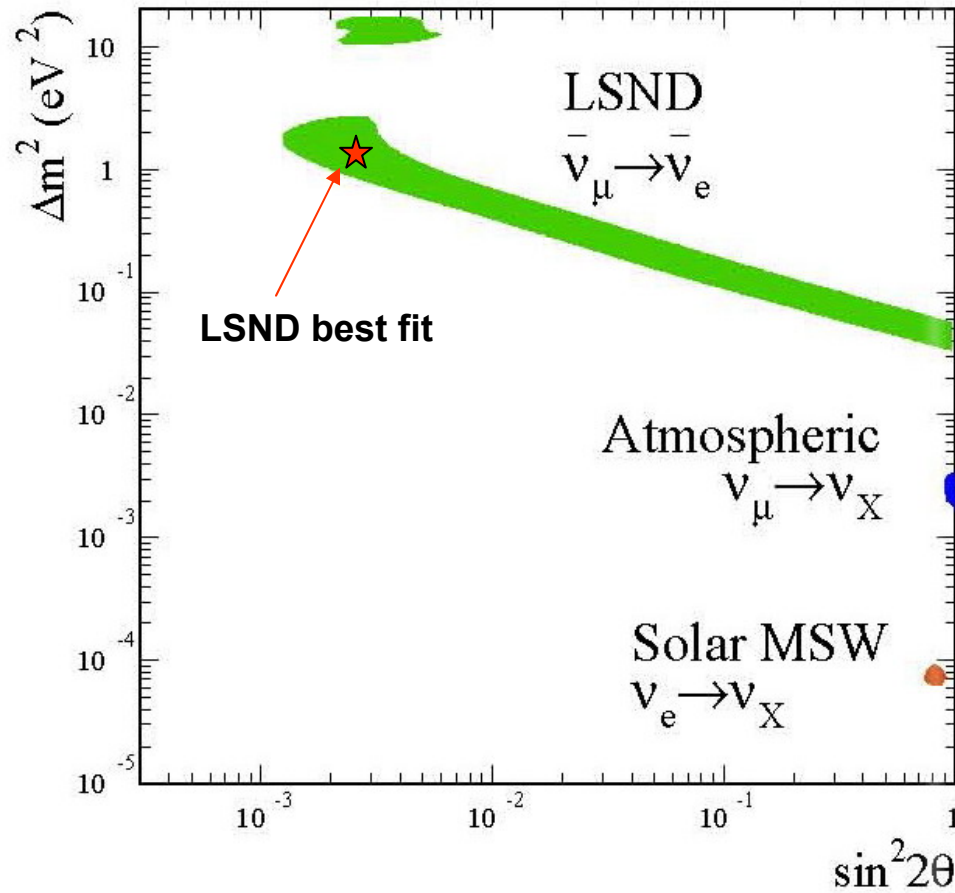
SK + K2K + MINOS

$$\theta_{13} \leq 10^\circ$$

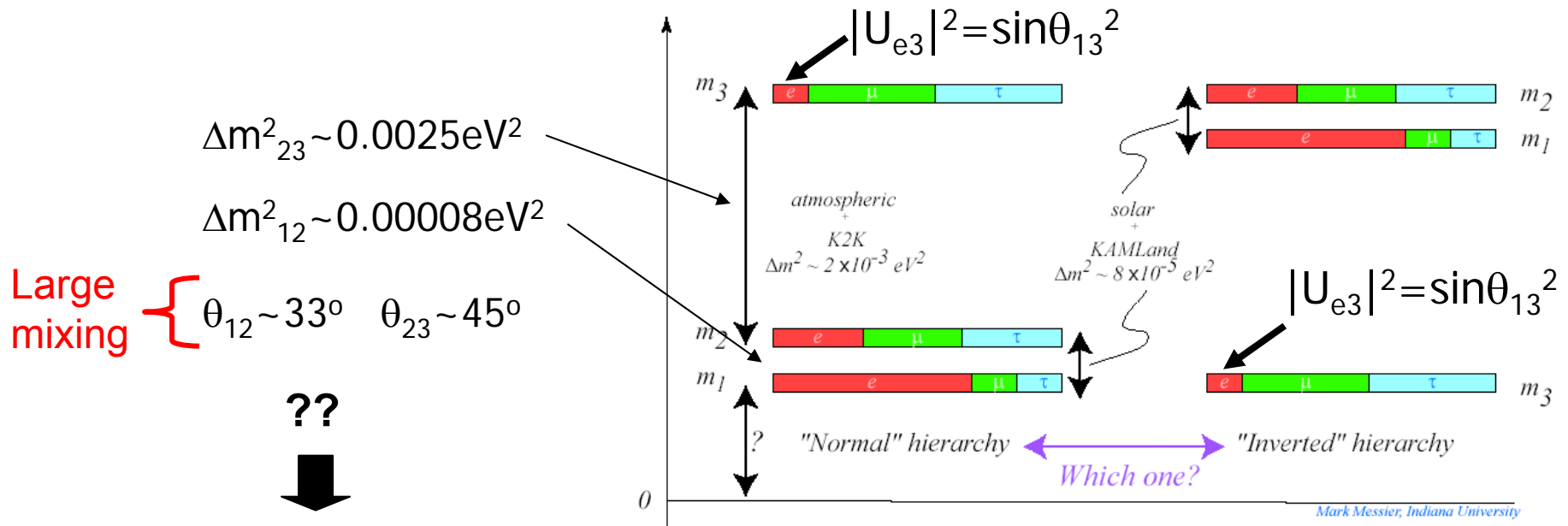


Only upper limit on  $\theta_{13}$   
No information about  $\delta$

# Oscillations before first MiniBooNe result

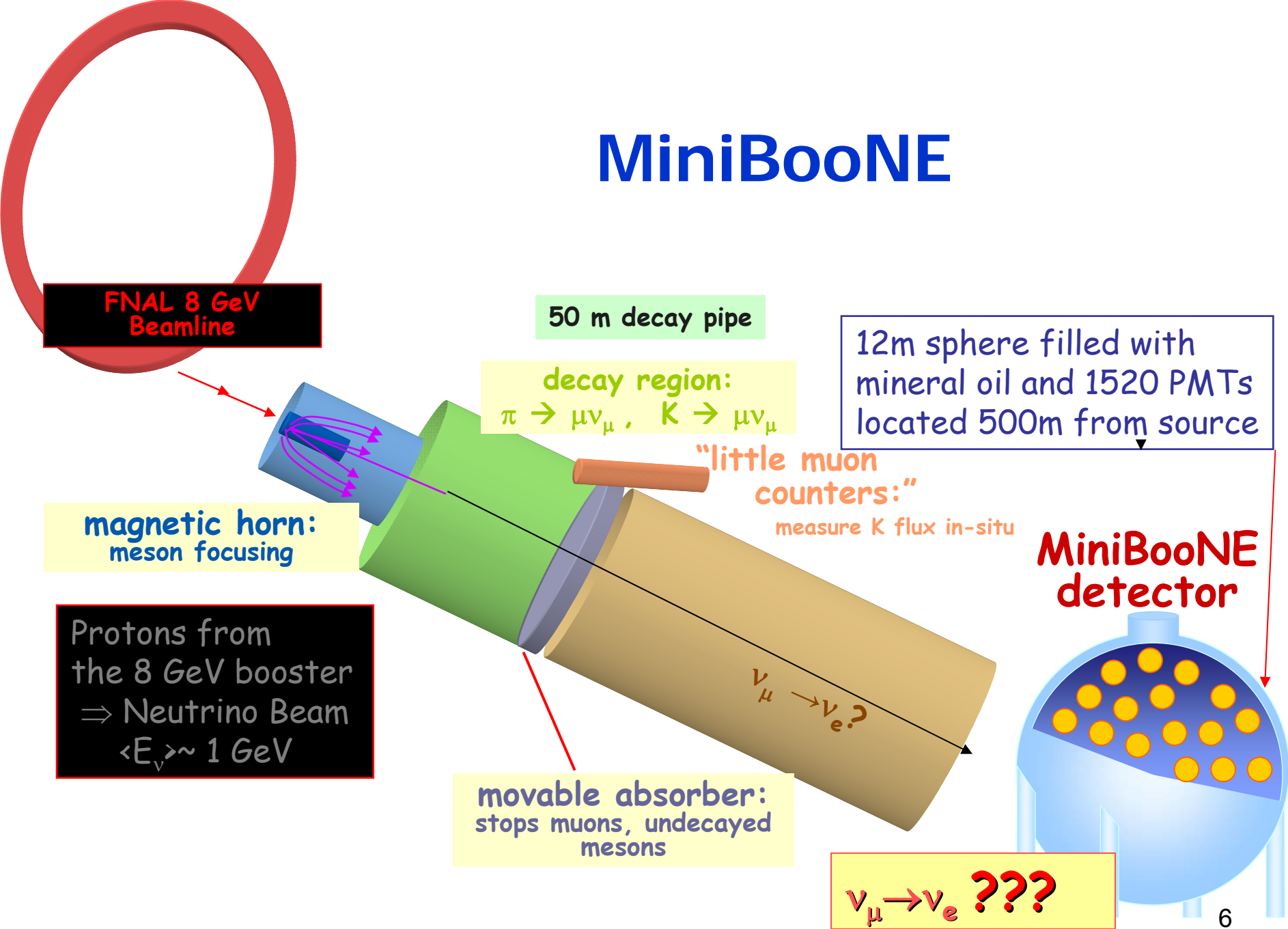


# Present knowledge and next steps



- Mixing angle  $\theta_{13}$
- Mass hierarchy (sign of  $\Delta m_{23}^2 \rightarrow m_3 > m_1$  or  $m_3 < m_1$ )
- CP violation
- Absolute mass scale
- Dirac or Majorana
- Approaches
  - LBL experiments: multi purpose ( $\theta_{13}$ ,  $\text{sign}(\Delta m^2)$ , CPV,  $\theta_{23}$ ,  $\Delta m_{23}^2$ )
  - Reactor-based  $\nu_e$  disappearance: single purpose ( $\theta_{13}$ ), complementary
  - Accelerator SBL  $\rightarrow$  LSND/MiniBooNE anomalies, sterile neutrinos (?)

# MiniBooNE



# MiniBooNE results

Neutrino run  
2007

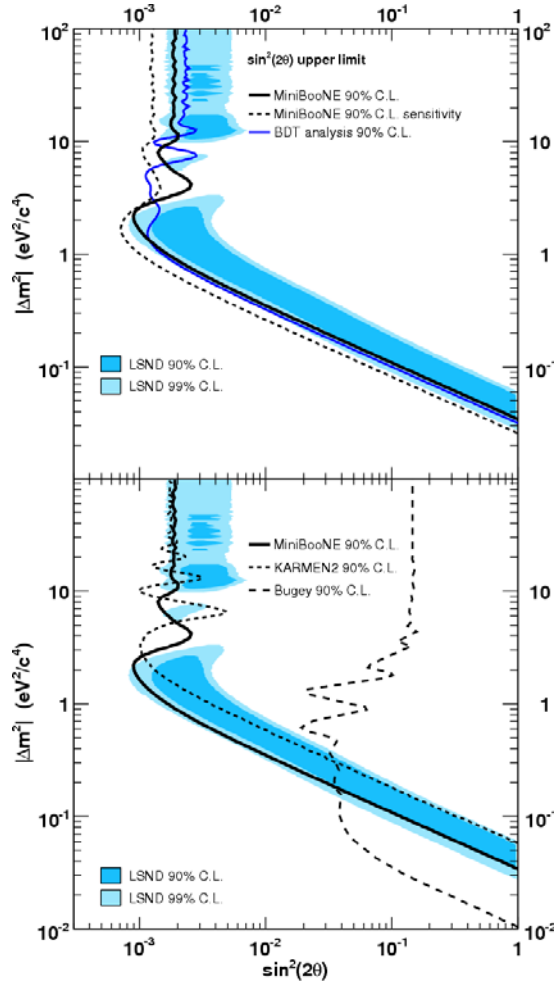
$\nu_\mu \rightarrow \nu_e$

$E_\nu > 475$  MeV

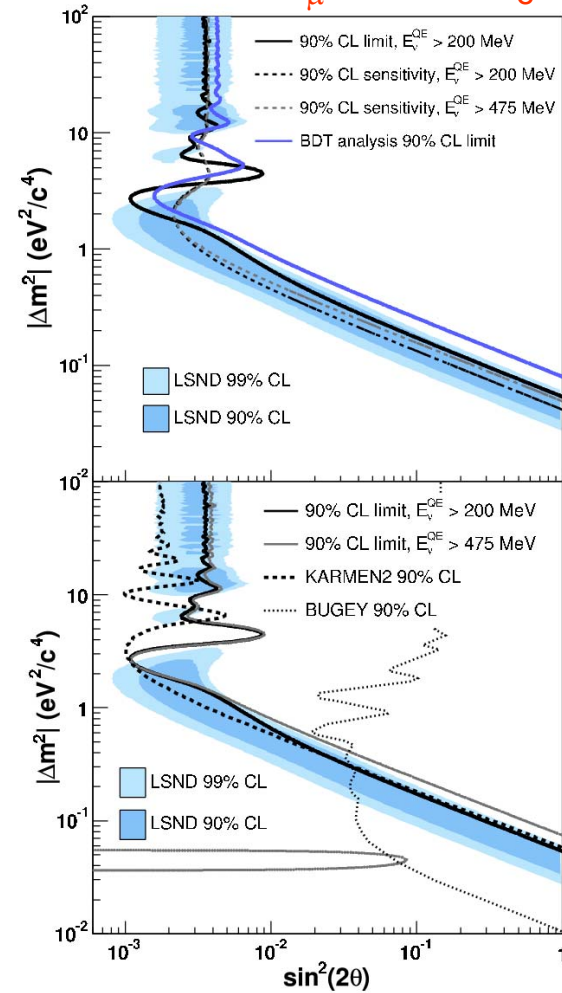
Antineutrino run

2009

$\text{anti-}\nu_\mu \rightarrow \text{anti-}\nu_e$



$5.58 \times 10^{20}$  POT

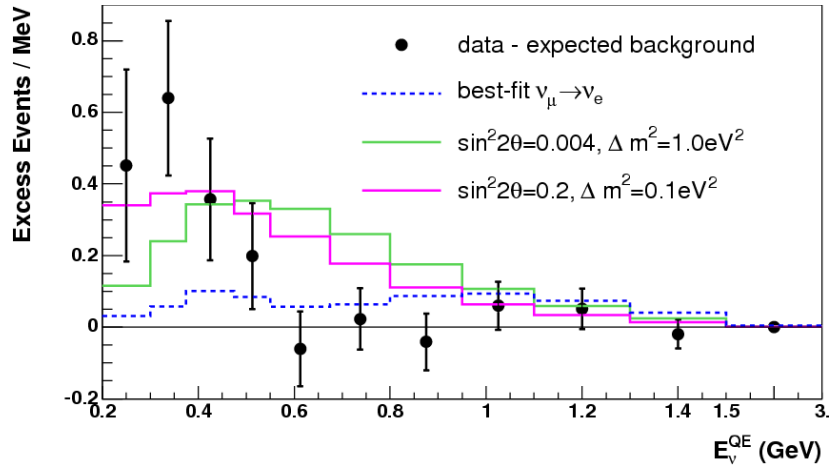
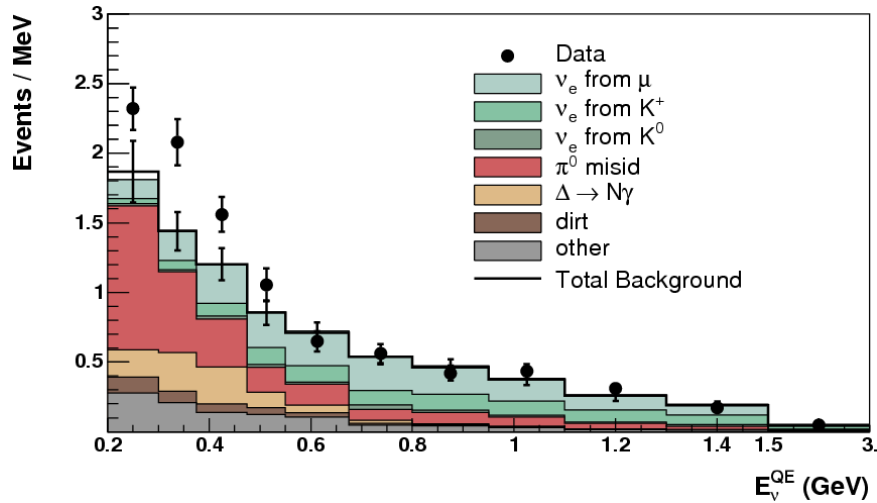


$3.39 \times 10^{20}$  POT

# MiniBooNE anomaly

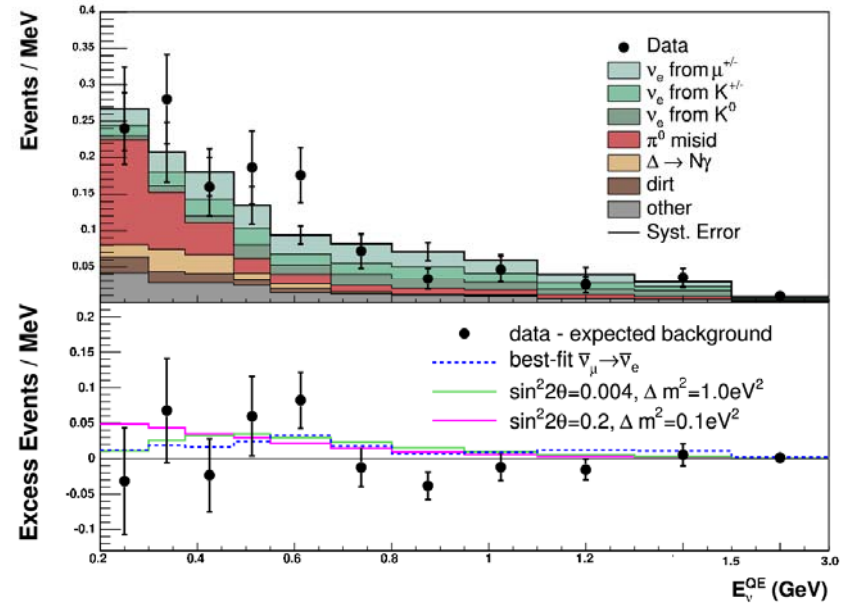
$$\nu_{\mu} \rightarrow \nu_e$$

arXiv:0812.2243 [hep-ex]



$$\text{anti-}\nu_{\mu} \rightarrow \text{anti-}\nu_e$$

arXiv:0904.1958 [hep-ex]





# MiniBooNE anomaly

Event Sample

anti- $\nu_e$  analysis  
( $3.39 \times 10^{20}$  POT)

$\nu_e$  analysis  
( $6.46 \times 10^{20}$  POT)

200 – 475 MeV

Data

61

544

Background

$61.5 \pm 11.7$

$415.2 \pm 43.4$

Excess

$-0.5 \pm 11.7$  ( $-0.04\sigma$ )

$128.8 \pm 43.4$  ( $3.0\sigma$ )

475 – 1250 MeV

Data

61

408

Background

$57.8 \pm 10.0$

$385.9 \pm 35.7$

Excess

$3.2 \pm 10.0$  ( $0.3\sigma$ )

$22.1 \pm 35.7$  ( $0.6\sigma$ )

# Possible explanations

## **Anomaly Mediated Neutrino-Photon Interactions at Finite Baryon Density**

**Jeffrey A. Harvey, Christopher T. Hill, & Richard J. Hill, arXiv:0708.1281**

## **CP-Violation 3+2 Model**

**Maltoni& Schwetz, arXiv:0705.0107; T. Goldman, G. J. Stephenson Jr., B. H. J. McKellar, Phys. Rev. D75 (2007) 091301**

## **Lorentz Violation**

**Katori, Kostelecky, & Tayloe, Phys. Rev. D74 (2006) 105009**

## **CPT Violation 3+1 Model**

**Barger, Marfatia, & Whisnant, Phys. Lett. B576 (2003) 303**

## **Heavy Sterile Neutrino Decay**

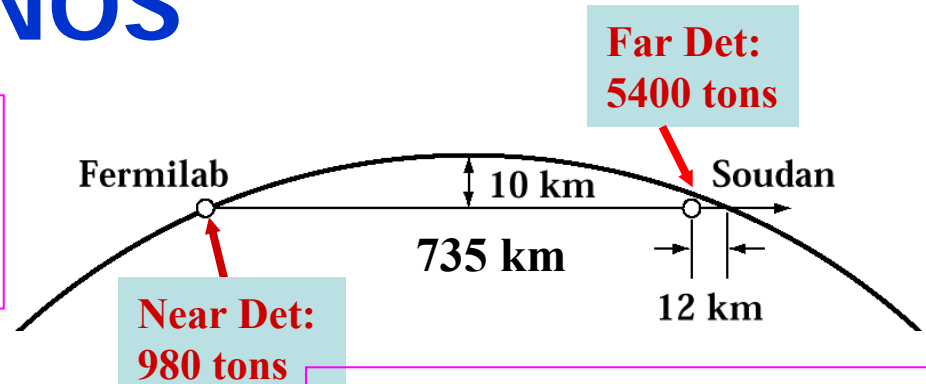
**S.N. Gninenko, arXiv:0902.3802**

## **VSBL Electron Neutrino Disappearance**

**Carlo Giunti& Marco Laveder, arXiv: 0902:1992**

# MINOS

Precise study of “atmospheric” neutrino oscillations, using the NUMI beam and two detectors

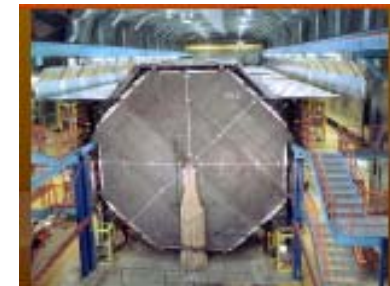
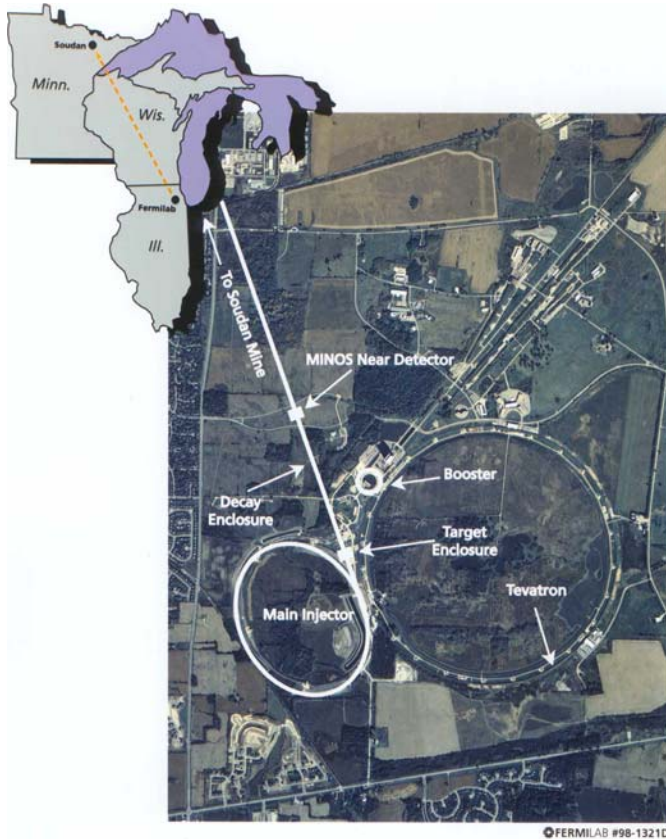


**Beam:** NuMI beam, 120 GeV  
Protons  $\rightarrow$   $\nu_{\mu}$  - beam

**Detectors:** ND, FD

**Far Det:** 5.4 kton magnetized Fe/Sci Tracker/Calorimeter at Soudan, MN (L=735 km)

**Near Det:** 980 ton version of FD, at FNAL (L  $\approx$  1 km)



# MINOS: $\nu_\mu \rightarrow \nu_s$

MINOS, PRL 101:221804,2008

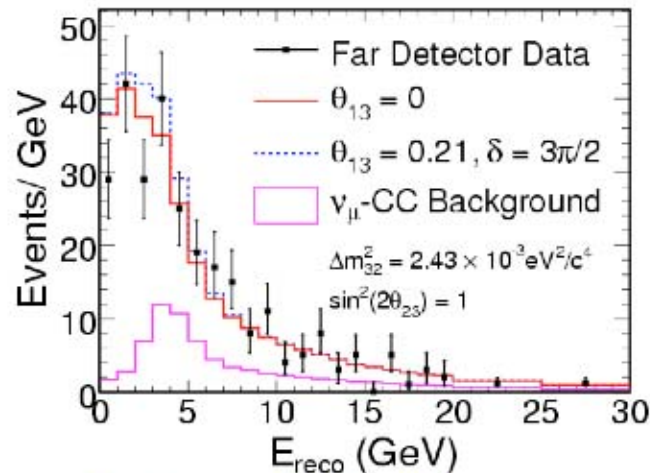
## Neutral Current Analysis

- **General NC analysis overview:**
  - All active neutrino flavours participate in NC interaction
  - Mixing to a sterile- $\nu$  will cause a deficit of NC events in Far Det.
  - Assume one sterile neutrino and that mixing between  $\nu_\mu$ ,  $\nu_s$  and  $\nu_\tau$  occurs at a single  $\Delta m^2$
- Survival and sterile oscillation probabilities become:

$$P(\nu_\mu \rightarrow \nu_\mu) = 1 - \alpha_\mu \sin^2(1.27 \Delta m^2 L / E)$$

$$P(\nu_\mu \rightarrow \nu_s) = \alpha_s \sin^2(1.27 \Delta m^2 L / E)$$

- ( $\alpha_{\mu,s}$  = mixing fractions)

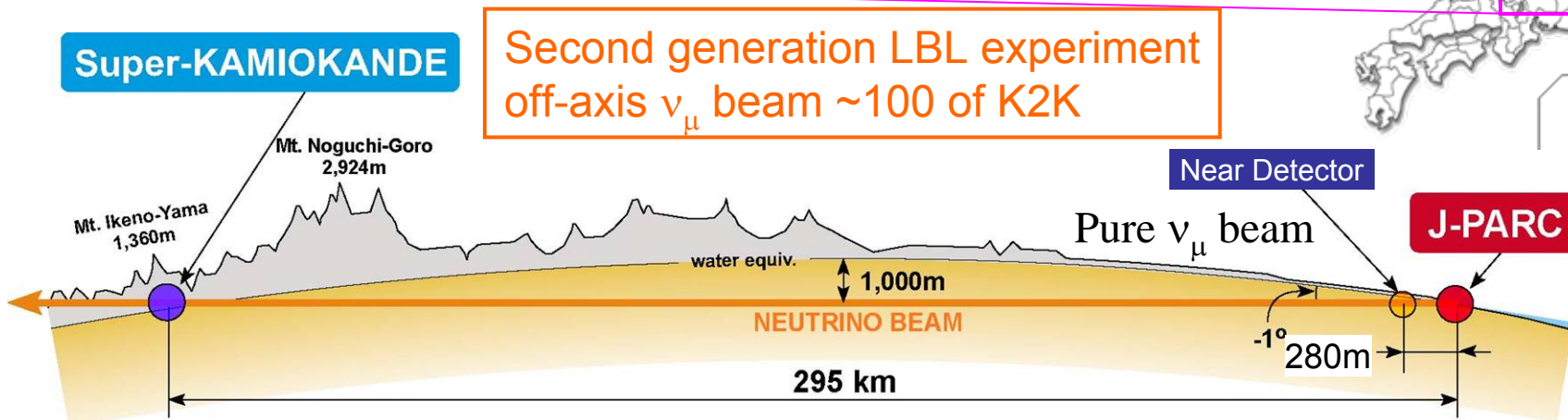
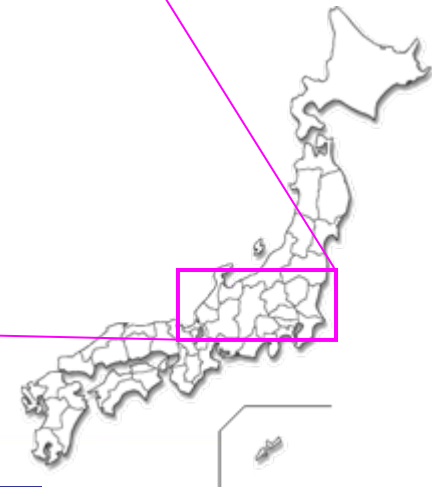
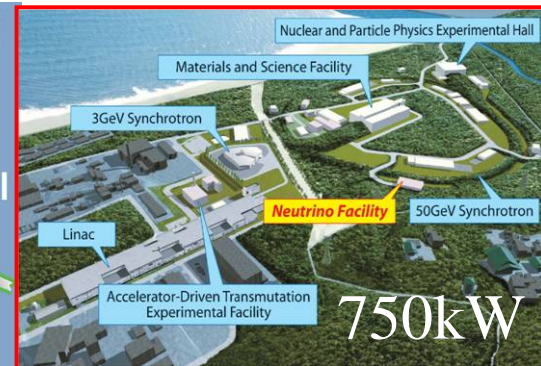
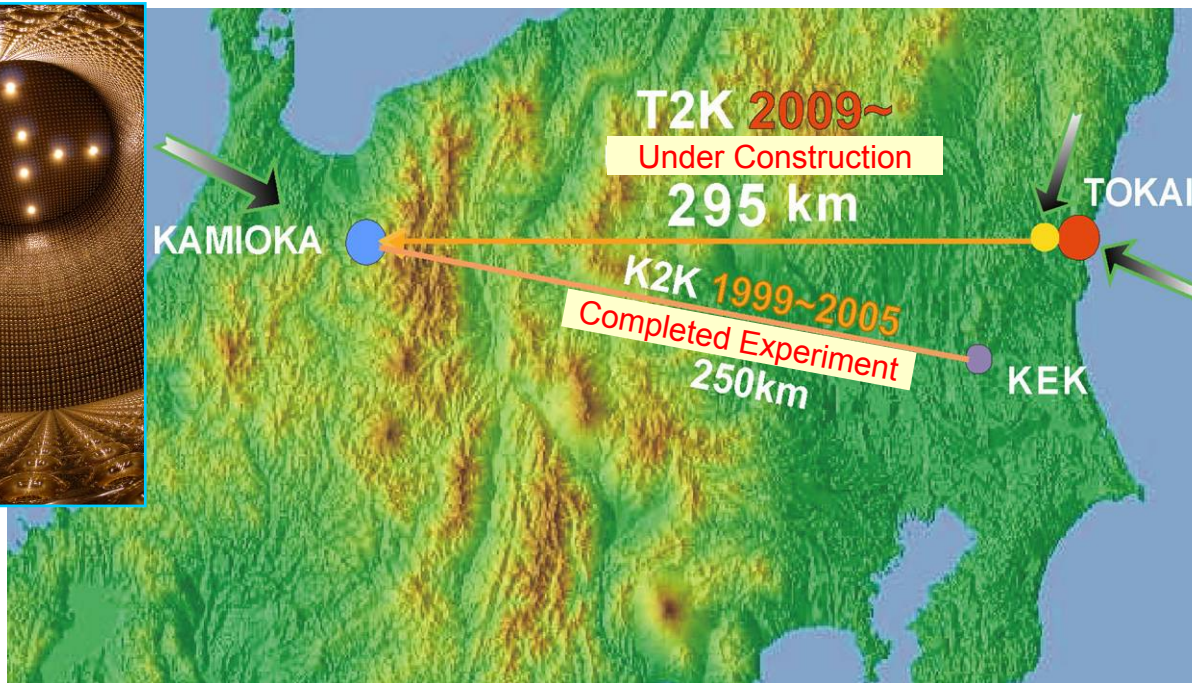
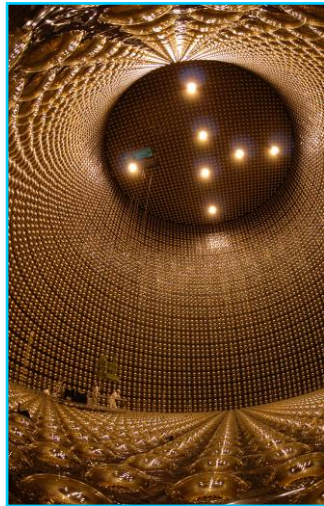


Simultaneous fit to CC and NC energy spectra yields the fraction of  $\nu_\mu$  that oscillate to  $\nu_s$ :

$$f_s = \frac{P(\nu_\mu \rightarrow \nu_s)}{1 - P(\nu_\mu \rightarrow \nu_\mu)} = 0.28_{-0.28}^{+0.25} (\text{stat.+syst.})$$

$$f_s < 0.68 \quad (90\% \text{ C.L.})$$

# T2K (Tokai to Kamioka) LBL neutrino experiment



# T2K

SuperKamiokande



KEK

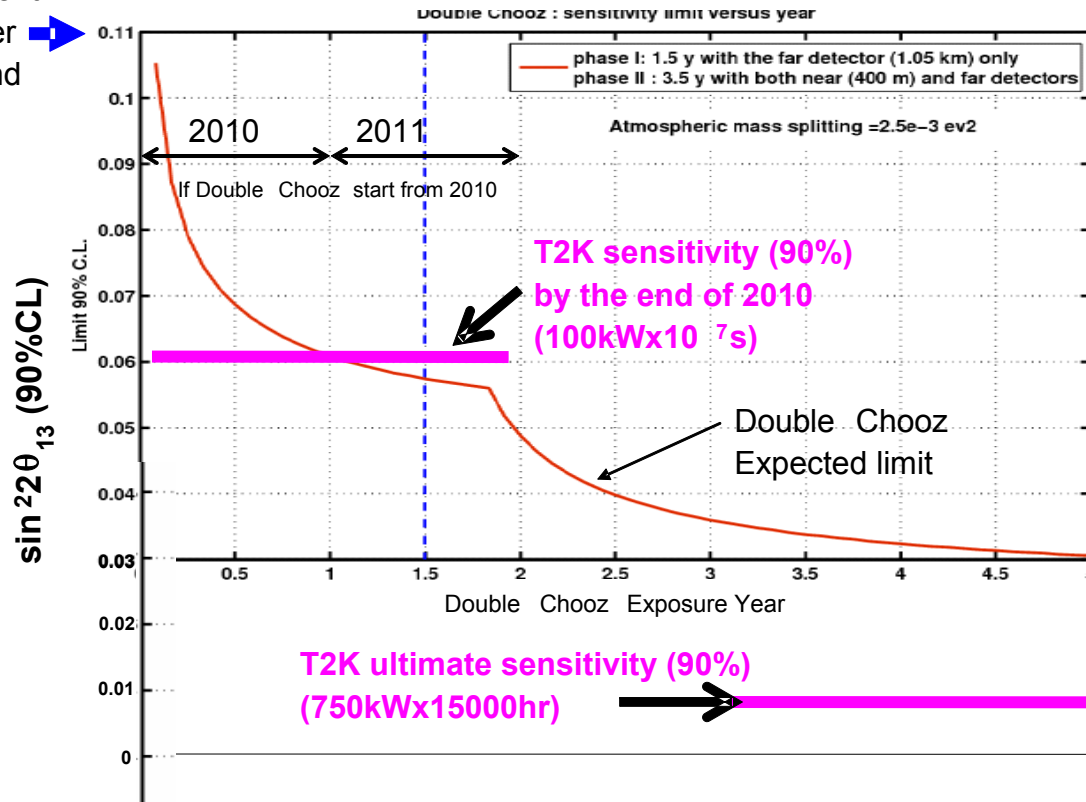
JPARC



# Expected sensitivity to $\theta_{13}$

Chooz

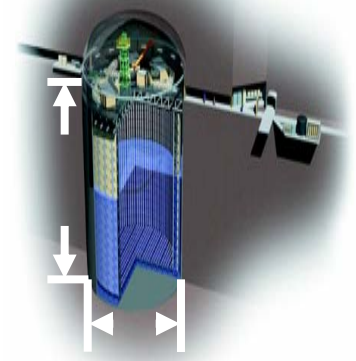
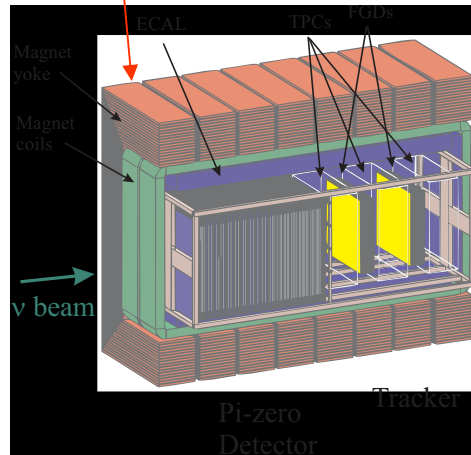
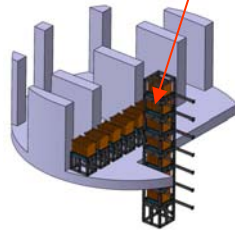
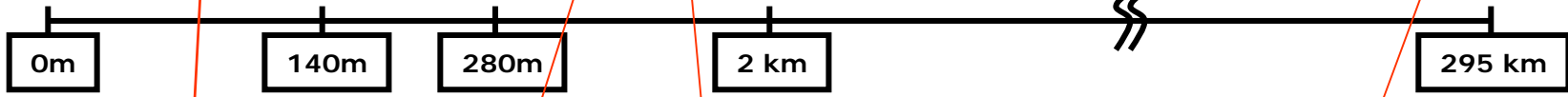
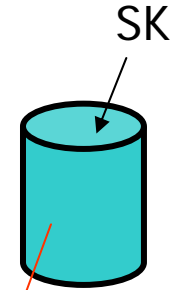
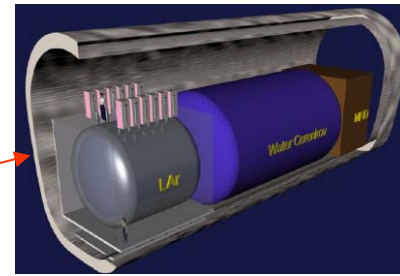
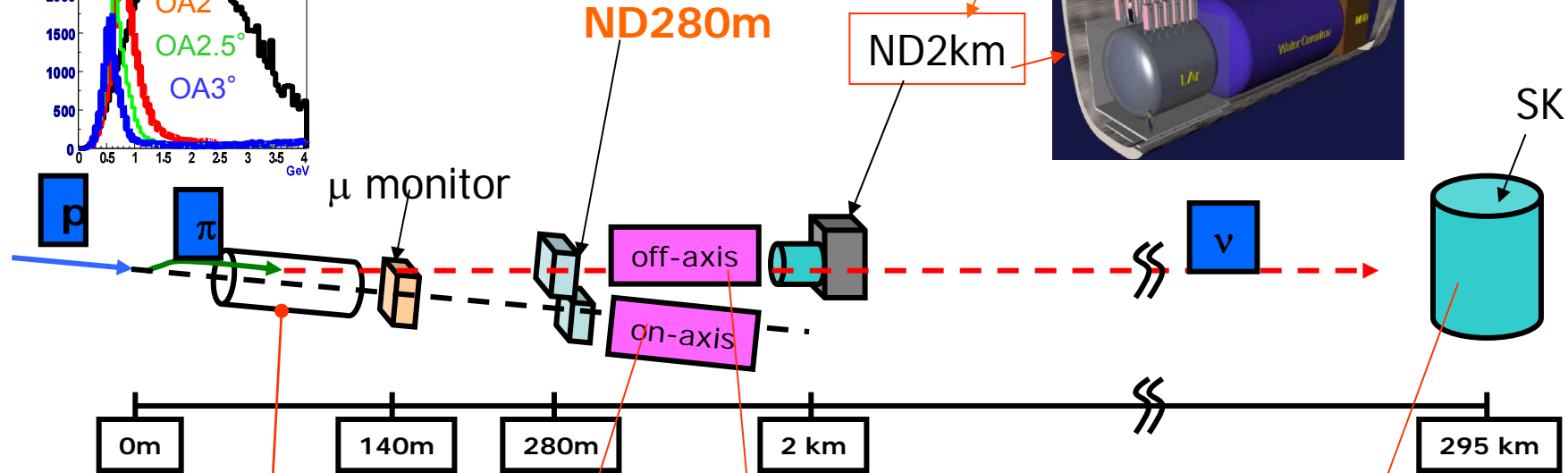
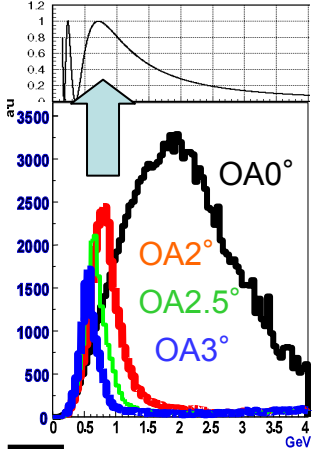
Present  
Upper Bound



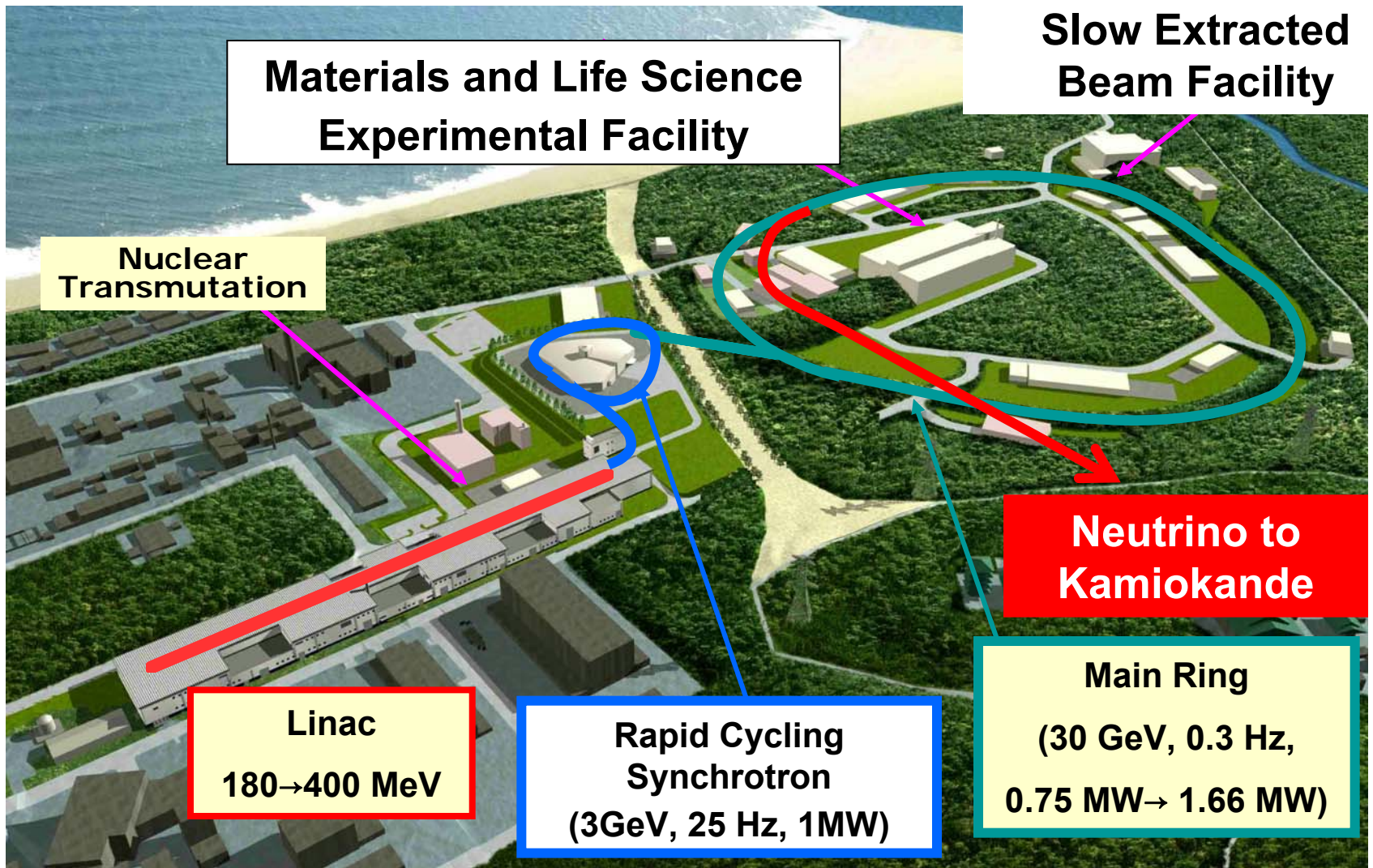


# T2K setup

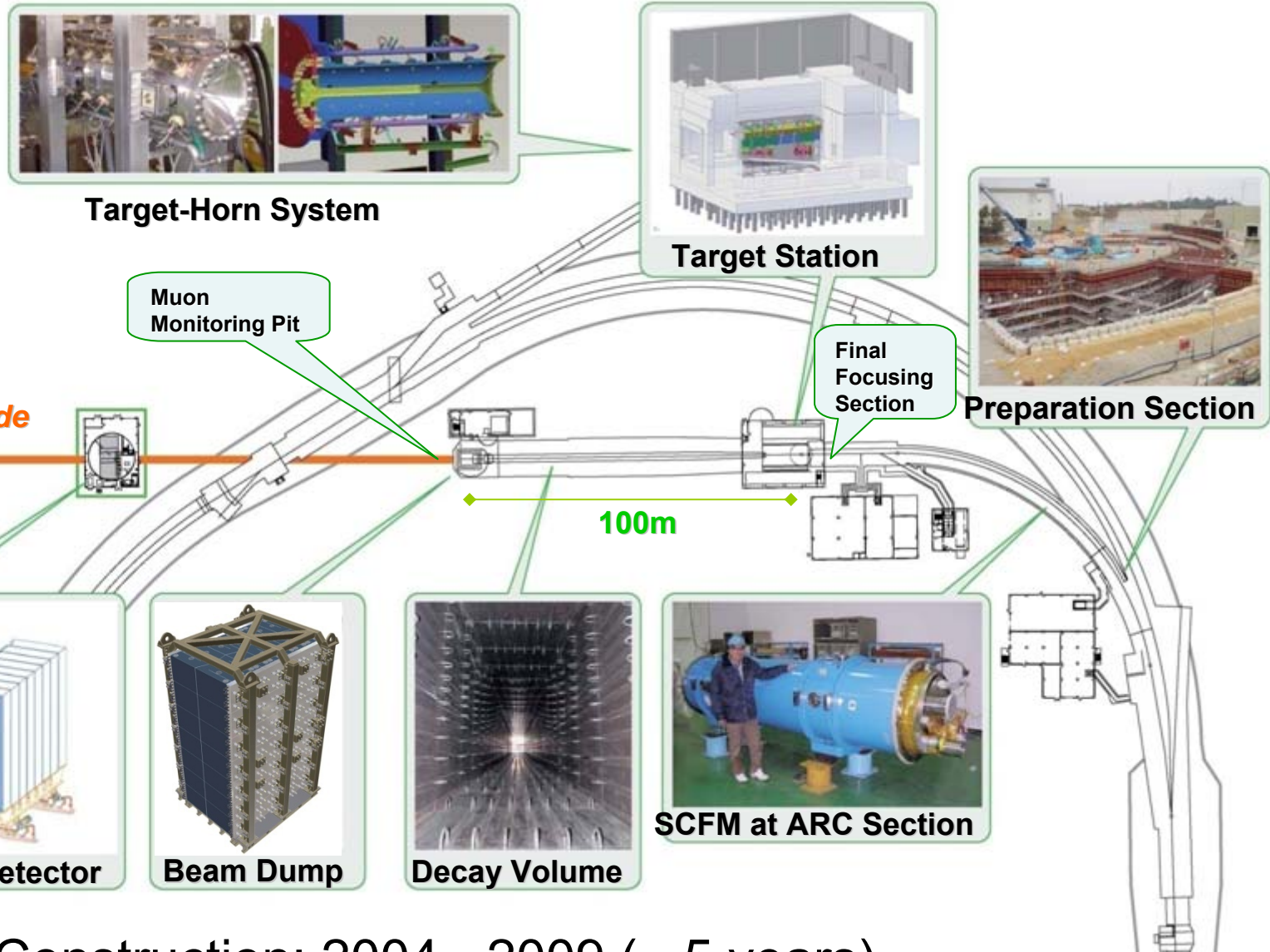
Possible Future



# J-PARC

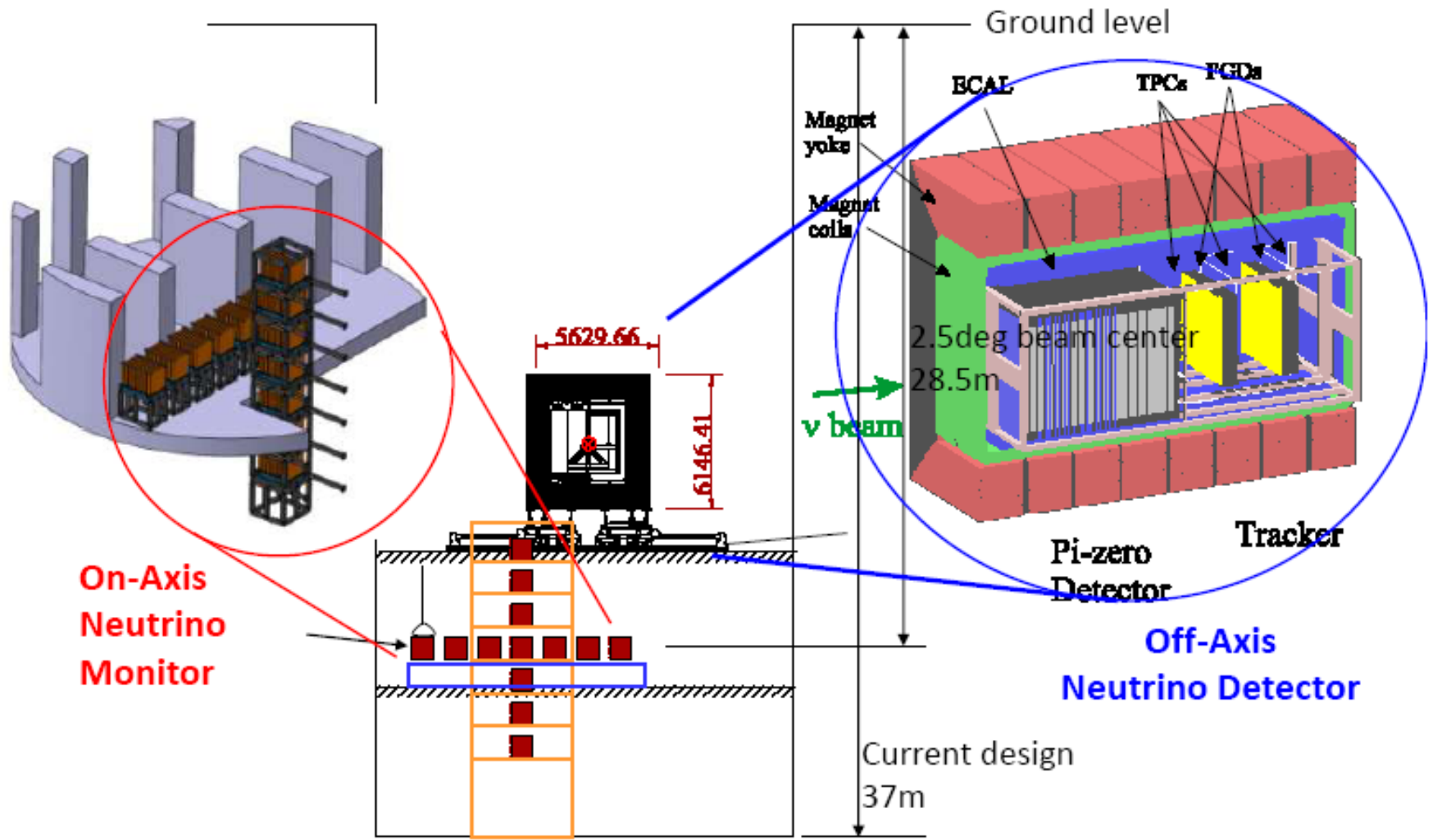


# Neutrino BeamLine

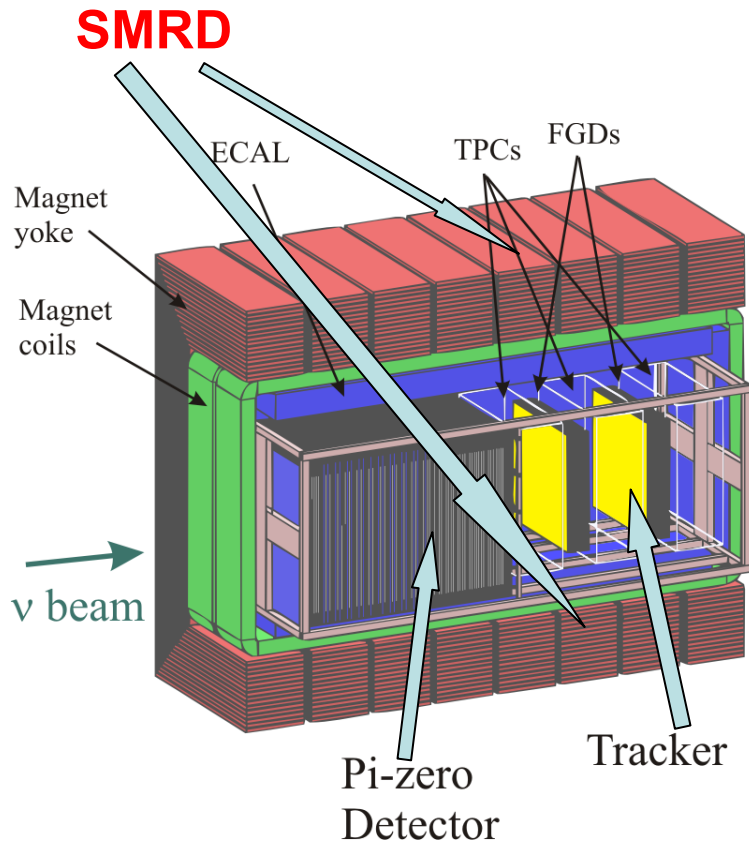


Construction: 2004 - 2009 (~ 5 years)

# Near Neutrino Detectors



# ND280 Off-axis Detector



**280m downstream from  
pion production target**

## UA1/NOMAD CERN magnet

operated at 0.2 T magnetic field

Tracker: Optimized for CC interactions measurements

- Fine Grained Detector (FGD)

- measure  $\nu$  beam flux,  $E_\nu$  spectrum, flavor composition through CC  $\nu$ -interactions,
- backgrounds CC- $1\pi$
- water and scintillator target

- Time Projection Chamber (TPC)

- measure charged particle momenta, particle ID via  $dE/dx$
- measure backgrounds/pion cross section

## Pi-Zero Detector (P0D)

- Optimized for NC  $\pi^0$  measurement
- measure  $\nu_e$  contamination

## Electromagnetic Calorimeter (ECAL)

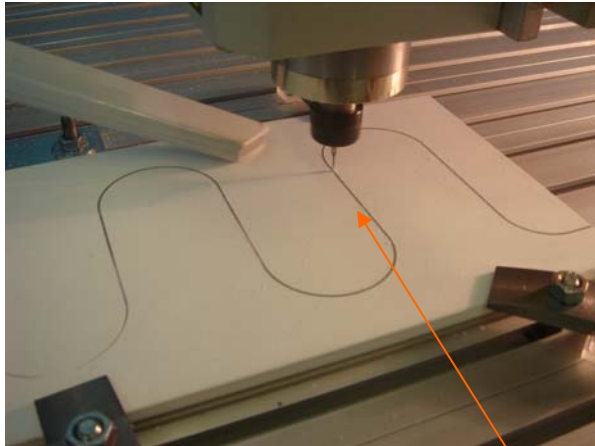
- Photon detection (from  $\pi^0$ ) in P0D and tracker
- charge particle ID and reconstruction

## Side Muon Range Detector (SMRD)

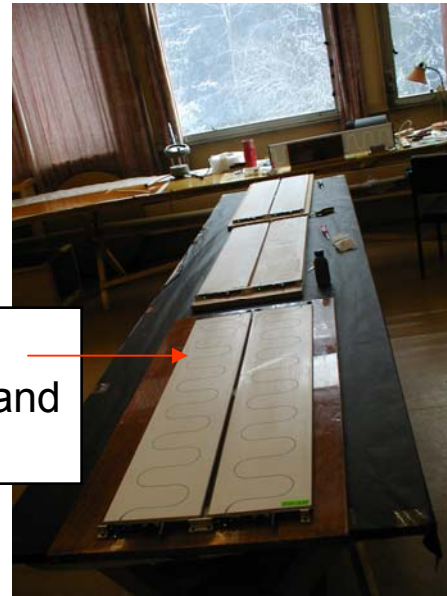
- measure momentum for lateral muons
- cosmic rays trigger

# SMRD detectors

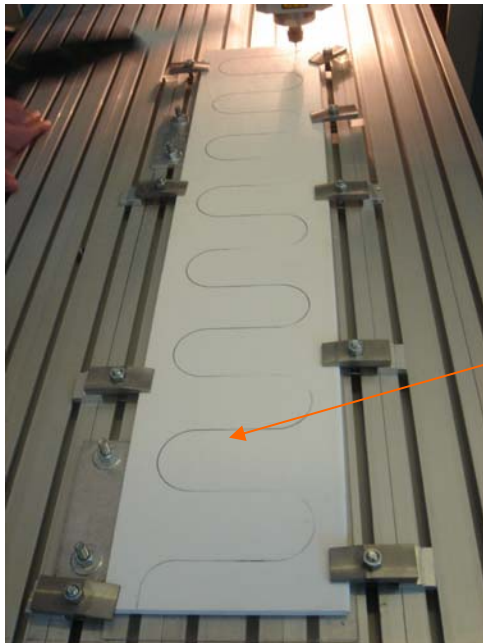
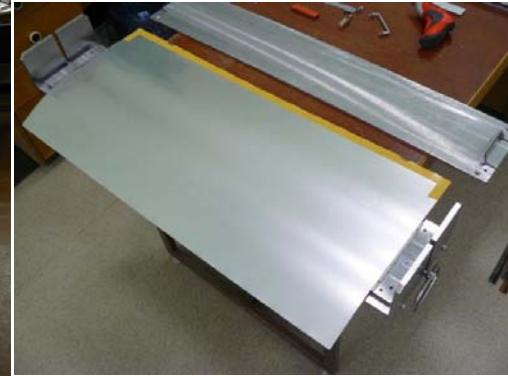
**INR  
Workshop**



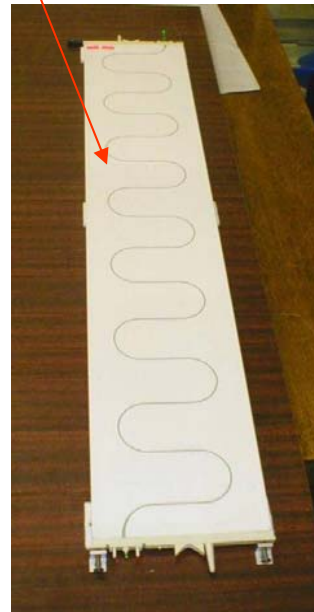
Y11 fibers  
embedded and  
glued



stainless steel  
container



Preparation  
of S-grooves



Ready for shipment

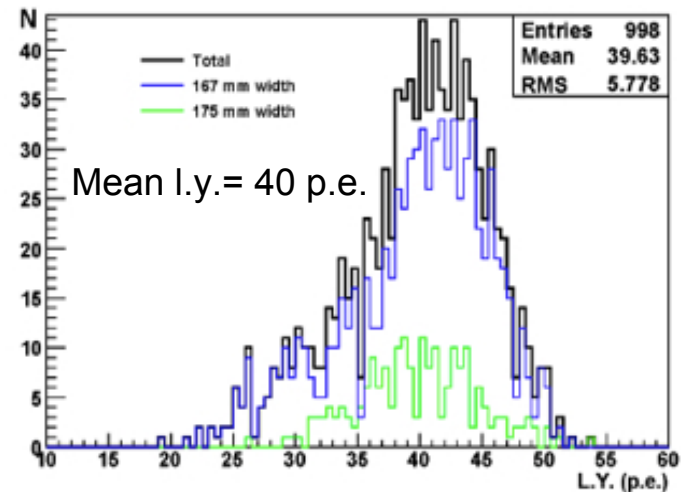


# 2130 SMRD detectors are manufactured at INR in 2007-2009



Completed February 2009

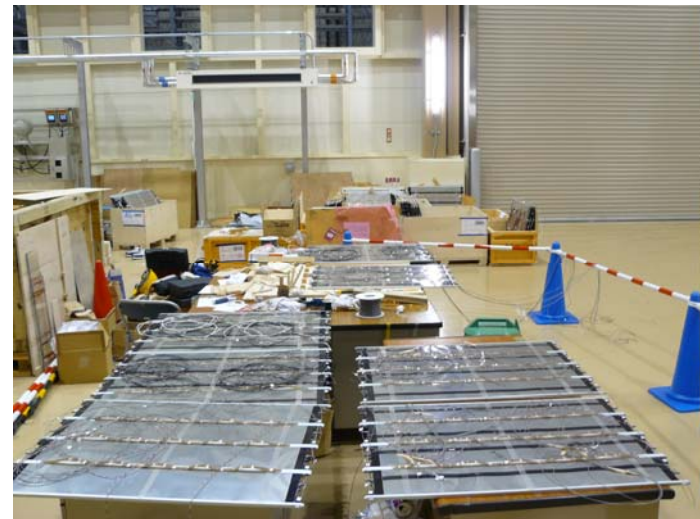
Shipped to JPARC in March 2009



Assembly at JPARC



# Assembly of SMRD modules at JPARC





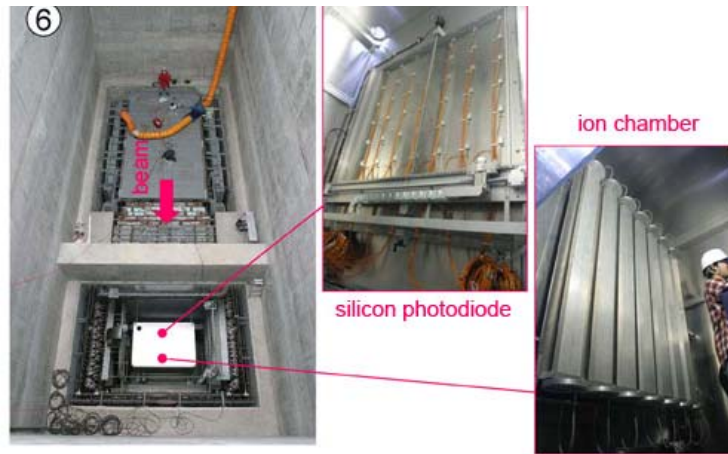
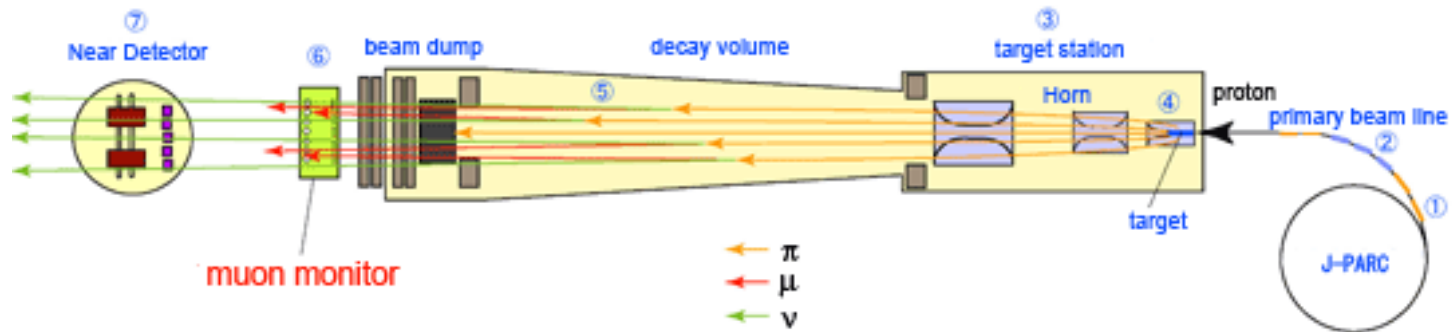
# Detectors in UA1 magnet



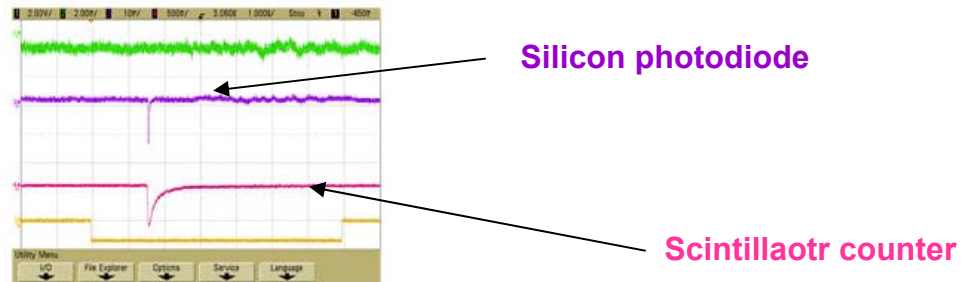
SMRD module (4 detectors) installed into magnet yoke

About 40% installed by 1 May 2009

# First $\nu$ 's for T2K



Muon signals  
24 April 2009



# T2K Physics run

Beam commissioning: April-May 2009, Detector completion: Fall 2009

Data taking start      December 2009

100kW, 30 GeV,  $10^7$  sec

$\nu_\mu \rightarrow \nu_e$

3.7 events at CHOOZ limit

background 0.25 ( $\nu_\mu$  NC ) 0.39 (beam  $\nu_e$ )

$\nu_\mu \rightarrow \nu_\mu$

(FCFV  $\mu$  -like)

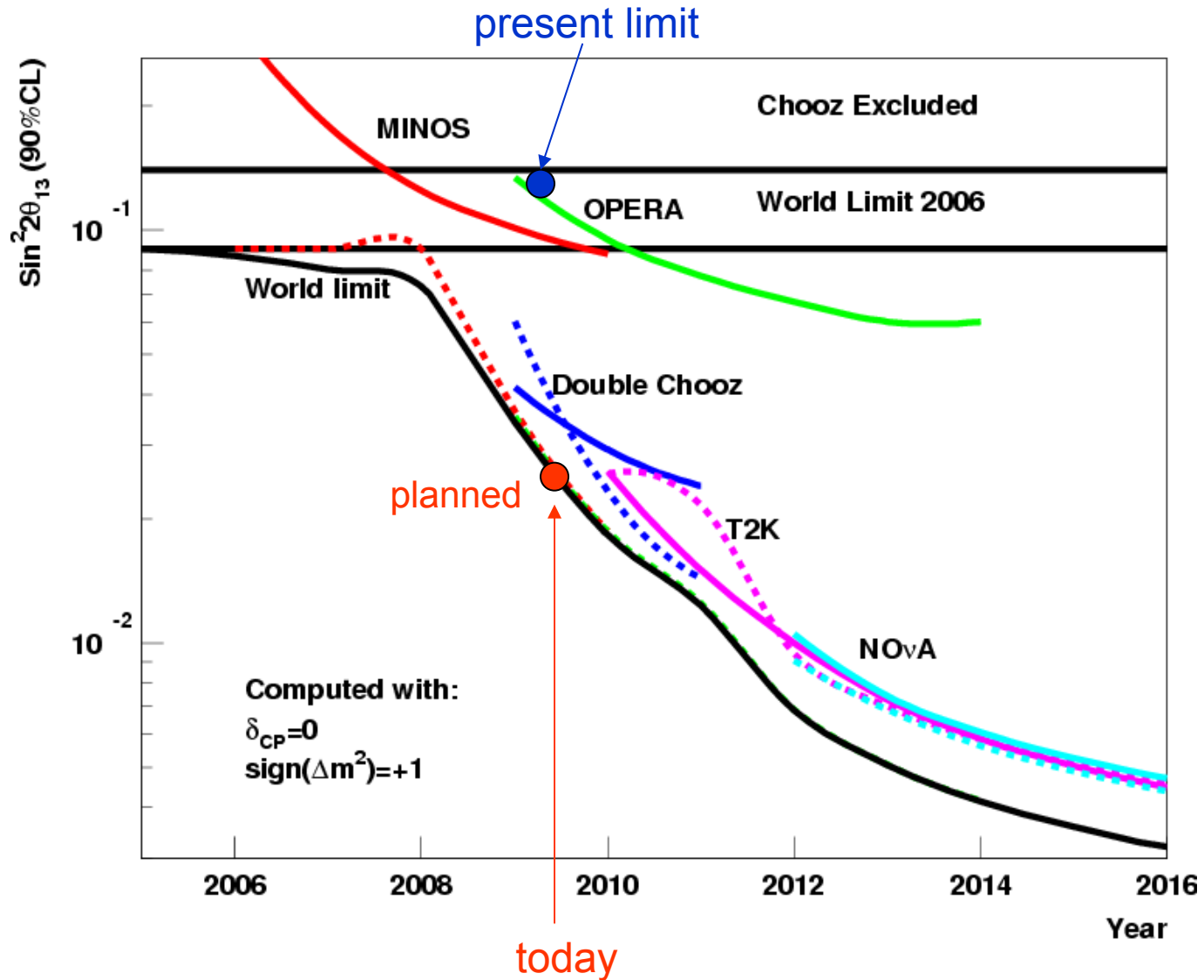
oscillation  
parameters

	null oscillation	oscillation
All	183.2	64.4
CCQE	118.0	22.9
CC non-QE	58.7	35.1
NC	6.5	6.5

$\sin^2 2\theta_{23} = 1.0$   
 $\Delta m^2_{23} = 2.4 \times 10^{-3} \text{ eV}^2$   
 $L = 295 \text{ km}$

# PREDICTIONS FOR SENSITIVITY TO $\theta_{13}$

A.Blondel et al. hep-ex/0606111



# Conclusion

**MINOS, OPERA**

data taking

**MiniBooNe**

**new anomaly appears**  
run with anti- $\nu$  beam

**T2K-I**

**first neutrino beam** in April 2009  
start data taking in December 2009

**Nova**

finally approved  
construction begins in May 2009

**MicroBooNE**  
**OscSNS ORNL**

proposal  
proposal



**LSND, MiniBooNE**  
**anomalies, sterile  $\nu$**